

## HW9 Solutions

### Chapter 10

#### Prob 12.

$$T = 960K \Rightarrow kT \simeq 0.0827 eV$$

$$\text{So, } \frac{N_2}{N_1} = \frac{d_2 e^{-E_2/kT}}{d_1 e^{-E_1/kT}} = 3 e^{-(E_2 - E_1)/kT} \simeq 0.146$$

With,  $N = N_1 + N_2$ , we can get  $N_1 \simeq 0.87N$ ,  $N_2 \simeq 0.13N$ .

#### Prob 14.

$v_p$  occurs at  $N(v)$  maximum. So,

$$\begin{aligned} \frac{dN}{dv} = 0 &\Rightarrow 4\pi N \left(\frac{m}{2\pi kT}\right)^{3/2} [2v + v^2(-2mv/2kT)] e^{-mv^2/2kT} = 0 \\ &\Leftrightarrow 2v - \frac{mv^3}{kT} = 0 \\ &\Leftrightarrow v_p = \sqrt{\frac{2kT}{m}} \end{aligned}$$

#### Prob 15.

$$T = 293K \Rightarrow kT \simeq 0.02525 eV$$

$$(a) E_m = \frac{3}{2} kT \simeq 0.0379 eV$$

$$(b) dN = N(E_m) dE = \frac{2N_A}{\sqrt{\pi} (kT)^{3/2}} \sqrt{E_m} e^{-E_m/kT} dE \simeq 2.78 \times 10^{21}$$

#### Prob 25.

$$E_F = 3.00 eV$$

$$(a) T = 295K \Rightarrow kT \simeq 0.0254 eV$$

$$\text{So, } \frac{N(E) dE}{V} = \frac{8\sqrt{2}\pi m^{3/2}}{h^3} \sqrt{E} \frac{1}{e^{(E-E_F)/kT} + 1} dE \simeq 1.03 \times 10^{-7} m^{-3}$$

$$(b) T = 2500K \Rightarrow kT \simeq 0.215 eV$$

$$\text{So, } \frac{N(E) dE}{V} = \frac{8\sqrt{2}\pi m^{3/2}}{h^3} \sqrt{E} \frac{1}{e^{(E-E_F)/kT} + 1} dE \simeq 1.48 \times 10^{23} m^{-3}$$

**Prob 27.**

(a) Mass of star  $M = N \times m_n$  ( $m_n$  : mass of neutron)

$$E_{grav} = -\frac{3}{5} \frac{GM^2}{R} = -\frac{3}{5} \frac{GN^2 m_n^2}{R}$$

$$E_{neut} = NE_m = \frac{3}{5} NE_F = \frac{3}{5} N \frac{h^2}{2m_n} \left( \frac{3N}{8\pi V} \right)^{2/3}$$

$$E = E_{grav} + E_{neut}$$

Minimum radius occurs at  $dE/dR = 0$ , so, we can get the radius solving this :

$$R = \frac{h^2}{GN^{1/3} m_n^3} \left( \frac{9}{32\pi^2} \right)^{2/3}$$

(b)  $M_{sun} \simeq 2.00 \times 10^{30} \text{ kg}$ , so  $M = 3M_{sun} \simeq 6.00 \times 10^{30} \text{ kg}$

$$\& N = \frac{M}{m_n} \simeq 3.6 \times 10^{57}$$

So,  $R \simeq 8.6 \text{ km}$

(c) Density  $\rho = \frac{M}{V} \simeq 2.2 \times 10^{18} \text{ kg/m}^3$

**Prob 28.**

(a)  $M_{sun} \simeq 2.00 \times 10^{30} \text{ kg}$ , so  $M = 2M_{sun} \simeq 4.00 \times 10^{30} \text{ kg}$

$$\& N = \frac{M}{m_n} \simeq 2.4 \times 10^{57}$$

Fermi energy

$$E_F = \frac{h^2}{2m_n} \left( \frac{3N}{8\pi V} \right)^{2/3} \simeq 140 \text{ MeV}$$

This value is small compared to the rest energy of the neutron ( $940 \text{ MeV}$ ), so we can use classical kinematics.

(b) The de Broglie wavelength

$$\lambda = \frac{h}{p} = \frac{h}{\sqrt{2mE}} \simeq 2.4 \text{ fm}$$

Also, from  $R = \frac{h^2}{GN^{1/3} m_n^3} \left( \frac{9}{32\pi^2} \right)^{2/3} \simeq 9.8 \text{ km}$ , we can get neutron number density

$$\frac{N}{V} \simeq 6.1 \times 10^{44} \text{ m}^{-3}$$

So, average neutron spacing is  $\left( \frac{N}{V} \right)^{-1/3} \simeq 1.18 \text{ fm}$

de Broglie wavelength is about two times larger than average neutron spacing.